

heit folgt. Die Voraussetzung Gl. (3) ist die schwächste heute bekannte Forderung für die Ergodizität, jedoch ist nicht bewiesen, daß sie optimal ist.

Zusammenfassend ergibt sich also, daß man als Bedingungen für ergodisches Verhalten an das Spektrum

die Forderung $\bar{\nu}$ nicht groß gegen 1 und an die ergodischen Observablen die Forderung (3) stellt.

Herrn Professor Dr. LUDWIG danke ich für zahlreiche Diskussionen und Anregungen zum Ergodenproblem.

The Motion of Mercury according to the Theory of Thiry and Lichnerowicz

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The extended gravitational theory of THIRY and LICHNEROWICZ¹ is based on the equation

$$G_b^a + \varrho v^a v_b = 0, \quad (1)$$

where G_b^a is EINSTEIN's tensor in five dimensions, ϱ a scalar function of matter and v^a a certain "five-velocity". In four dimensions (1) yields EINSTEIN's 10 and MAXWELL's 4 equations, all of them modified by a 15th field function

$$g_{55} = 2 e^{2\sigma}, \quad (2)$$

which is determined by a 15th field equation. For the static field outside a spherically symmetric body without an electric charge only 4 of those equations are linearly independent. These yield

$$\left. \begin{aligned} (e^\sigma)'' &= 2, & (\sigma' e^\sigma)' &= (\nu' e^\sigma)' = 0, \\ 2\mu'' + \mu'^2 + \nu'^2 + 3\sigma'^2 &= 0, \end{aligned} \right\} \quad (3)$$

if the four-dimensional metric is² written:

$$\left. \begin{aligned} ds^2 &= e^{-\sigma} (e^{-\nu} dr^2 + e^{\mu-\nu} d\omega^2 - e^\nu dt^2) \\ \text{with } d\omega^2 &= d\beta^2 + \cos^2 \beta d\varphi^2. \end{aligned} \right\} \quad (4)$$

A rigorous solutions of (3) is

$$\left. \begin{aligned} \nu &= \frac{m}{a} \ln \frac{r-a}{r+a}, & \sigma &= \frac{n}{a} \ln \frac{r-a}{r+a}, \\ \mu &= \ln(r^2 - a^2), & a^2 &= m^2 + \frac{3}{4} n^2, \end{aligned} \right\} \quad (5)$$

with the constants m, n uniquely determined by the central body. Applying the well known formalism of super-potentials to the³ extended theory, we receive

$$m = (\varrho), \quad 6n = 2m - (e^{2\sigma} F_{\mu\nu} F^{\mu\nu}), \quad (6)$$

where the brackets mean a multiplication by $e^\sigma/8\pi$ and the three-dimensional integration over the field-generating body. If the MAXWELL field $F_{\mu\nu}$ inside this body would be purely radiational, then $F_{\mu\nu} F^{\mu\nu} = \mathfrak{B}^2 - \mathfrak{E}^2 = 0$; but inside the sun we must also assume a COULOMB con-

tribution. Thus we have even

$$F_{\mu\nu} F^{\mu\nu} < 0, \quad \text{hence } n > \frac{1}{3} m. \quad (7)$$

Now it is easily⁴ seen, that the five-velocity v^a of a testbody always describes a five-dimensional geodesic. In the special case of an electrically neutral test-body this also means a geodesic four-velocity (in general this is modified² not only by the LORENTZ force, but also by an additional term dependent on the variability of σ). For a motion in the plane $\beta=0$ of the metric (4) this yields (with $D, E = \text{const}$):

$$(dr/d\varphi)^2 = (E^2 e^{2\mu-2\nu} - e^{-\nu-\sigma}) D^{-2}, \quad (8)$$

giving with (5) in the usual approximation a nearly circular ellipse with a slowly moving perihelion. The angular velocity of the latter turns out² to have EINSTEIN's value multiplied by the factor

$$f = \frac{6m+n}{6m-3n} > \frac{19}{15} > \frac{5}{4}, \quad (9)$$

where the numerical estimate followed from (7). Thus we have a relative deviation above $\frac{1}{4} = 25\%$, whereas the observations of mercury⁵ have confirmed EINSTEIN's expression with an uncertainty below 2%.

It must be emphasized that our result has nothing to do with the question, whether the new field variable (2) shall be called a variable⁶ gravitational "constant", a dielectricity⁷ of the vacuum or something else. To show this independence of an interpretation of σ clearly, we have not given it any name at all.

Implicitly the above result is already contained in former⁸ papers, where we have considered JORDAN's extended theory of gravitation. There the situation could be saved by a certain condition on the undetermined parameters of the field equations. But the theory of THIRY and LICHNEROWICZ¹ — which was independently developed also by SCHMUTZER⁷ — is based on the uniquely arranged equation (1). Thus being a completely determined special case of JORDAN's⁶ theory, its contradiction with experience cannot be avoided.

The detailed calculations have been performed in 1955 within a research for the Deutsche Forschungsgemeinschaft under Prof. LUDWIG.

¹ Y. THIRY, J. Math. Pures Appl. 30, 275 [1951]; A. LICHNEROWICZ, Theories Relativistes de la Gravitation et de l'Électromagnétisme, Masson & Cie., Paris 1955, p. 153–179.

² K. JUST, Zur JORDANSchen Gravitations-Theorie 1957 (unpublished).

³ K. JUST, Z. Phys. 142, 493 [1955].

⁴ A. LICHNEROWICZ, l. c. ¹, p. 214.

⁵ G. M. CLEMENCE, Rev. Mod. Phys. 19, 361 [1947].

⁶ P. JORDAN, Schwerkraft und Weltall, Vieweg, Braunschweig 1955.

⁷ E. SCHMUTZER, Z. Phys. 149, 329 [1957]; A. LICHNEROWICZ, l. c. ¹, p. 199–201.

⁸ K. JUST, Z. Phys. 144, 411 [1956]. K. JUST, Unified Field Theories, 1956 (unpublished). That the unusual, but wholly general formulation (4) of a static metric with spherical symmetry gives for (1) the exact solution (5), was discovered² later.

